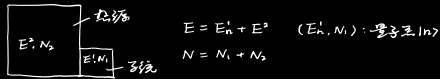


• 巨正则子系

宏观状态: (T, V, μ) . 允许能量, 粒子数的交换



巨正则状态数:

$$\Omega(E, N, V) = \sum_{E_h, N_1} \Omega_1(E_h, N_1) \cdot \Omega_2(E - E_h, N - N_1)$$

对 $\ln \Omega_2$ 展开:

$$\begin{aligned} \ln \Omega_2(E - E_h, N - N_1) &\simeq \ln \Omega_2(E, N) - \frac{\frac{\partial \ln \Omega_2}{\partial E} E_h}{\frac{1}{k_B T}} - \frac{\frac{\partial \ln \Omega_2}{\partial N} N_1}{-\frac{\mu}{k_B T}} \\ &= \ln \Omega_2(E, N) - \beta E_h + \beta \mu N_1. \end{aligned}$$

代入:

$$\Omega(E, N, V) = \sum_{E_h, N_1} \Omega_1(E_h, N_1) \cdot e^{-\beta(E_h - \mu N_1)}$$

• 根据微正则子系的概率分布, 得到:

$$p(E_h, N_1) = \frac{e^{-\beta(E_h - \mu N_1)}}{\sum_{E_h, N_1} e^{-\beta(E_h - \mu N_1)}} = \frac{1}{Q} e^{-\beta(E_h - \mu N_1)}$$

$$Q = \sum_{E_h, N_1} e^{-\beta(E_h - \mu N_1)}, \text{ 巨配分函数 (教科书中的 } \Xi \text{)}$$

注: 有的书上会定义 $\alpha = -\beta\mu$. $\Rightarrow e^{-\beta E_h - \alpha N}$. (T, μ) 是独立变量, 故可取 (α, β) 为独立变量.

后面会定义 $e^{\beta\mu} = z$, 逸度, fugacity.

回到巨配分函数

$$Q = \sum_N e^{\beta\mu N} \sum_{E_h} e^{-\beta E_h(N)} = \sum_N e^{\beta\mu N} \frac{\Omega(E_h, N)}{\text{正则配分函数}} \quad [\text{对比: } Z = \sum_E \Omega(E) e^{-\beta E}]$$

• [乍看之下回到了正则子系, 但这只是形式上的,

真正的简化是操作层面的]

与正则子系同理, 求和可用极值来代表:

$$\begin{aligned} Q &\simeq e^{\beta\mu N^*} \cdot Z(T, V, N^*) = e^{\beta\mu N^*} e^{-\beta F^*} \\ &= e^{-\beta(F^* - \mu N^*)} = e^{-\beta \bar{\Psi}^*} \end{aligned}$$

$$\bar{\Psi} = F - \mu N \text{ 巨热力学, } \bar{\Psi} = E - TS - \mu N = -pV.$$

$$\text{此处: } \bar{\Psi} = -k_B T \ln Q.$$

热力学量

$$\langle E \rangle = - \left(\frac{\partial}{\partial \beta} \ln Q \right)_\mu. \quad (\text{假设 } z = e^{\beta\mu} \text{ 变量, 否则全拿下来 } \mu N)$$

$$\langle N \rangle = \left(\frac{\partial}{\partial (\beta\mu)} \ln Q \right)_\beta.$$

由

$$d\bar{\Psi} = -SdT - pdV - Nd\mu.$$

• 可以求出其他的热力学函数. [$\bar{\Psi} = -pV$, 因 p 已经知道了]

• 理想气体

$$Z(T, V, N) = \frac{1}{N!} \left(\frac{V}{\lambda_T^3} \right)^N.$$

$$Q = \sum_{N=0}^{\infty} z^N \cdot \frac{1}{N!} \left(\frac{V}{\lambda_T^3} \right)^N = \exp \left(\frac{zV}{\lambda_T^3} \right).$$

$$\ln Q = \frac{zV}{\lambda_T^3}.$$

$$\bar{\Psi} = -k_B T \ln Q = -k_B T \frac{zV}{\lambda_T^3}.$$

$$p = -\frac{\bar{\Psi}}{V} = \frac{k_B T}{V} \frac{z}{\lambda_T^3}. \quad N = ?$$

$$\langle N \rangle = \frac{\partial}{\partial (\beta\mu)} \ln Q = \frac{V}{\lambda_T^3} \frac{\partial}{\partial (\beta\mu)} (e^{\beta\mu}) = \frac{zV}{\lambda_T^3}.$$

$$\Rightarrow p = \langle N \rangle k_B T.$$

证: 由 $N = \frac{2V}{\lambda_T^3}$: $e^{\beta\mu} = \frac{N}{V} \lambda_T^3 = n \lambda_T^3$.

$\mu = \frac{1}{\beta} \ln(n \lambda_T^3)$. 又得到了这个结果.

$\langle E \rangle = - \left(\frac{\partial}{\partial \beta} \ln \Omega \right)_z = -2V \frac{\partial}{\partial \beta} (\lambda_T^{-3})$.

$\lambda_T = \frac{h}{\sqrt{2\pi m k_B T}} = \sqrt{\frac{\beta}{2\pi m}} h$, $\frac{\partial}{\partial \beta} (\lambda_T^{-3}) = -\frac{3}{2} \lambda_T^{-3} \cdot \frac{1}{\beta}$

$\Rightarrow \langle E \rangle = \frac{3}{2} 2V \cdot \frac{1}{\beta \lambda_T^3} = \frac{3}{2} \langle N \rangle \cdot k_B T$.

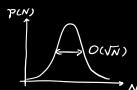
• 能量与粒子的涨落

$\langle N^2 \rangle_c = \left(\frac{\partial^2}{\partial (\beta\mu)^2} \ln \Omega \right)_\beta = \frac{\partial N}{\partial (\beta\mu)} \Big|_\beta = k_B T \left(\frac{\partial N}{\partial \mu} \right)_T \propto N$. ↗ 严格来说是 $\left(\frac{\partial N}{\partial \mu} \right)_{T,V}$

$\Rightarrow \frac{\langle N^2 \rangle_c}{N} \propto \frac{1}{\sqrt{N}}$. [问: 热力学极限怎么体现?]

$\langle N^2 \rangle_c \propto N$.

故 N 的分布也是极窄的高斯:



[严格来说, $P(N)$ 满足泊松分布, 但 $\langle N \rangle \rightarrow \infty$ 时与高斯一致]

问: $\left(\frac{\partial N}{\partial \mu} \right)_{T,V} = ?$ 想法: 与某个响应函数取到一处.

$d\mu = -SdT + Vdp$

$\left(\frac{\partial \mu}{\partial N} \right)_{T,V} = \left(\frac{\partial \mu}{\partial P} \right)_{T,V} \left(\frac{\partial P}{\partial N} \right)_{T,V} = \frac{V}{N} \left(\frac{\partial P}{\partial N} \right)_{T,V}$

又 $\left(\frac{\partial P}{\partial N} \right)_V \cdot \left(\frac{\partial N}{\partial V} \right)_P \left(\frac{\partial V}{\partial P} \right)_N = -1$ [T不变].

代入得

$\left(\frac{\partial \mu}{\partial N} \right)_{T,V} = - \left(\frac{V}{N} \right)^2 \left(\frac{\partial P}{\partial V} \right)_{N,T}$

$\Rightarrow \langle N^2 \rangle_c = -k_B T \cdot \left(\frac{V}{N} \right)^2 \left(\frac{\partial P}{\partial V} \right)_{N,T}$

$= k_B T \cdot \frac{N^2}{V^2} \underline{K_T} = n k_B T \cdot N K_T \propto N$.
是涨落

一般来讲, K_T 有限. 但在临界点附近, K_T 发散!!

这说明粒子数 N 的涨落是剧烈的! \rightarrow 临界乳光现象.

$\langle E^2 \rangle_c = \left(\frac{\partial^2}{\partial \beta^2} \ln \Omega \right)_z = - \left(\frac{\partial E}{\partial \beta} \right)_{z,N} = k_B T^2 \left(\frac{\partial E}{\partial T} \right)_{z,V}$

[注: $\left(\frac{\partial E}{\partial T} \right)_{z,V} \neq C_V$. 后者是 $\left(\frac{\partial E}{\partial T} \right)_{N,V}$]

结果: $\langle E^2 \rangle_c = \underbrace{k_B T^2 C_V}_{\text{有限项}} + k_B T \left(\frac{\partial E}{\partial N} \right)^2 \cdot \underbrace{\langle N^2 \rangle_c}_{\text{涨落项}}$.